FINAL PROGRESS REPORT FOR S-92504-Z:

A THREE-DIMENSIONAL ANALYSIS OF THE GALACTIC GAMMA-RAY EMISSION RESULTING FROM COSMIC-RAY INTERACTIONS WITH THE

INTERSTELLAR GAS AND RADIATION FIELDS

7N-90 1997 000 244

TASK LEADER: Thomas J. Sodroski

ATR: Eli Dwek

Description of Activities:

The contractor will provide support for the analysis of data under ADP (NRA 96-ADP-09; Proposal No . 167-96adp). The primary task objective is to contruct a 3-D model for the distribution of high-energy (20 MeV - 30 GeV) gamma-ray emission in the Galactic disk. Under this task the contractor will utilize data from the EGRET instrument on the Compton Gamma-Ray Observatory, H I and CO surveys, radio-continuum surveys at 408 MHz, 1420 MHz, 5 GHz, and 19 GHz, the COBE Diffuse Infrared Background Experiment (DIRBE) all-sky maps from 1 to 240 µm, and ground-based B, V, J, H, and K photometry. The respective contributions to the gamma-ray emission from cosmic ray/matter interactions, inverse Compton scattering, and extragalactic emission will be determined.

This task is supported by T. J. Sodroski (SAC).

Significant Accomplishments:

During this initial year of our 2-year program we have successfully completed the most critical part of our modelling effort, the derivation of a detailed 3-D map of the Galactic interstellar radiation field (ISRF) from 0.44 μm to 1000 μm based primarily on data from the DIRBE instrument on board NASA's Cosmic Background Explorer (COBE) satellite, and ground-based B, V, J, H, and K photometry. A radiative transfer code has been implemented to derive the radial and z-distance variations of the ISRF between 1 and 1000 µm. The code utilizes the stellar input parameters obtained from Freudenreich's (1996) analysis of the DIRBE 1 - 5 µm data, and the dust infrared spectra obtained from a recent 3-D analysis (Sodroski et al. 1997) of DIRBE 12 - 240 μm observations of the Galactic plane region. In order to account for

the processes of magnetic field and cosmic ray generation and diffusion in the Galaxy and the possible connection between a gamma-ray halo and the "thick radio disk".

- 3) The use of ISRF model to constrain the various models for dust grains in the interstellar medium.
- 4) Investigation of the effects of radiation pressure on the dynamics of dust grains in the interstellar medium.

References:

Arendt, R. G., et al. 1998, ApJ, in press

Arendt, R. G., Freudenreich, H. T., & Dwek, E. 1996, in AIP Conf. Proc. 348, Unveiling the Cosmic Infrared Background, ed. E. Dwek (New York: AIP), 305

Bloemen, J. G. B. M. 1985, A&A, 145, 391

Bronfman, L., Cohen, R. S., Alvarez, H., May, J., & Thaddeus, P. 1988, ApJ, 324, 248

Burton, W. B., & Gordon, M. A. 1978, A&A, 63, 7

Cohen, M. 1995, ApJ, 444, 874

Digel, S. 1991, Ph.D. thesis, Sterrewacht Leiden

Dwek, E., et al. 1995, ApJ, 445, 716

Dwek, E., et al. 1997, ApJ, 475, 565

Freudenreich, H. T. 1996, ApJ, 468, 663

Henderson, A. P., Jackson, P. D., & Kerr, F. J. 1982, ApJ, 263, 116

Jackson, P. D., & Kellman, S. A. 1974, ApJ, 190, 53

Knapp, G. R. 1988, in Proc. CITA Workshop, The Mass of Gas in the Galaxy, ed. M. Fich, p. 35

Mathis, J. S. 1990, ARA&A, 28, 37

Sodroski, T. J. 1988, Ph.D. thesis, University of Maryland

Sodroski, T. J., et al. 1997, ApJ, 480, 173

Wainscoat, R. J., et al. 1992, ApJS, 83, 111

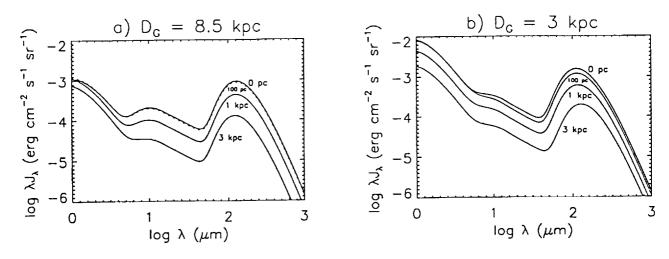


Fig. 1.—(a) Galactic interstellar radiation field at a Galactocentric distance of 8.5 kpc. Curves labeled 0 pc, 100 pc (dotted line), 1 kpc, and 3 kpc correspond to various z-distances above the Galactic midplane. (b) Same as (a), for a Galactocentric distance of 3 kpc.

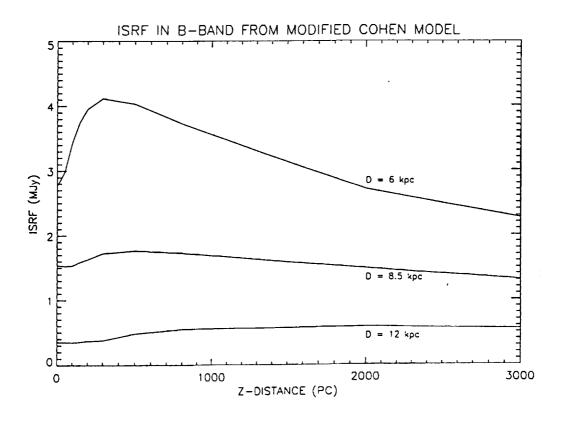


Fig. 2. — The z-distance variation of the Galactic interstellar radiation field in the B-band ($\lambda_{eff} = 0.44 \ \mu m$), for various Galactocentric distances D.

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Form Approved REPORT DOCUMENTATION PAGE OMB No. 0704-0188 Public reporting burden for this collection of information is estimated to average 1 hour per response, including the time for reviewing instructions, searching existing data sources, gathering and maintaining the data needed, and completing and reviewing the collection of information. Send comments regarding this burden estimate or any other aspect of this collection of information, including suggestions for reducing this burden, to Washington Headquarters Services, Directorate for Information Operations and Reports, 1215 Jefferson Davis Highway, Suite 1204, Arlington, VA 22202-4302, and to the Office of Management and Budget, Paperwork Reduction Project (0704-0188), Washington, DC 20503. 3. REPORT TYPE AND DATES COVERED 1. AGENCY USE ONLY (Leave blank) 2. REPORT DATE August 1998 Contractor Report 4. TITLE AND SUBTITLE 5. FUNDING NUMBERS A Three-Dimensional Analysis of the Galactic Gamma-Ray Emission Resulting From Cosmic-Ray Interaction With the Interstellar Gas and S-92504-Z Radiation Fields. 6. AUTHOR(S) Thomas J. Sodroski 8. PEFORMING ORGANIZATION 7. PERFORMING ORGANIZATION NAME(S) AND ADDRESS (ES) REPORT NUMBER 10. SPONSORING / MONITORING 9. SPONSORING / MONITORING AGENCY NAME(S) AND ADDRESS (ES) AGENCY REPORT NUMBER National Aeronautics and Space Administration CR-208638-1 Washington, DC 20546-0001 11. SUPPLEMENTARY NOTES 12a. DISTRIBUTION / AVAILABILITY STATEMENT 12b. DISTRIBUTION CODE Unclassified-Unlimited Subject Category: 90 Report available from the NASA Center for AeroSpace Information, 7121 Standard Drive, Hanover, MD 21076-1320. (301) 621-0390. 13. ABSTRACT (Maximum 200 words) This final report outlines the progress of the contractor's support for the analysis of data under ADP (NRA 96-ADP-08; Proposal No. 167-96adp). The primary task object was to construct a 3-D model for the distribution of high-energy (20 MeV-30GeV) gamma-ray emission in the Galactic disk. Under this task the contractor was to utilize data from the EGRET instrument on the Compton Gamma-Ray Observatory, HI and CO surveys, radio-continuum surveys at 408 MHz, 1420 MHz, 5 GHz, and 19 GHz, the COBE Diffuse Infrared Background Experiment (DIRBE) all-sky maps from 1 to 240 um,

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